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# EM Propagation Models and Scintillation Effects

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LANL  
23 July 2019

## Homogeneous Shell

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- Definite top and bottom altitudes
- No variation in electron density or magnetic field
  - (1) Taylor expansion of refractive index ( $f \gg f_p, f_{ce}$ ) with and without magnetic birefringence - LOS: all rays reach satellite.
  - (2) Two dimensional Snell's law in spherical geometry (Bourger's) - solve analytically for rays that reach satellite at each frequency.

## Non-Homogeneous Shell (2D)

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- Can have bottom and top altitudes, or not.
- Electron density and magnetic field vary with altitude ( $r$ ) only – no gradients in other directions
  - (3) Two dimensional Snell's law in spherical geometry (Bourger's) – solve analytically for rays that reach satellite at each frequency, including  $n_e(r)$  and  $B_o(r)$  in calculation.
  - (4) Ray trace – WKB approximation (geometric optics). must iterate over frequency and angle.
  - (5) Divide ionosphere into spherical shells of constant  $n_e$  and  $B_o$ . Match boundary conditions at each shell interface.

## Non-Homogeneous Shell (3D)

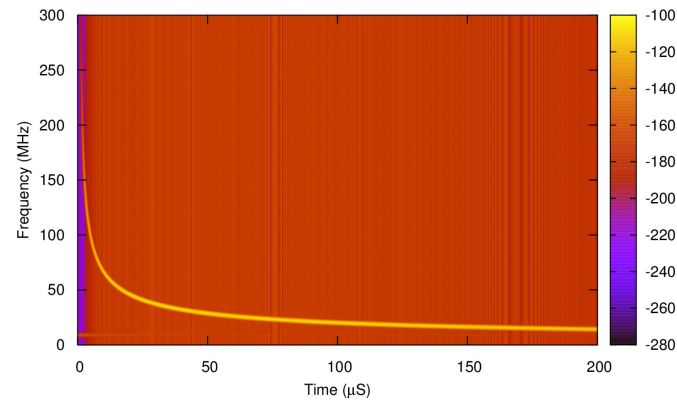
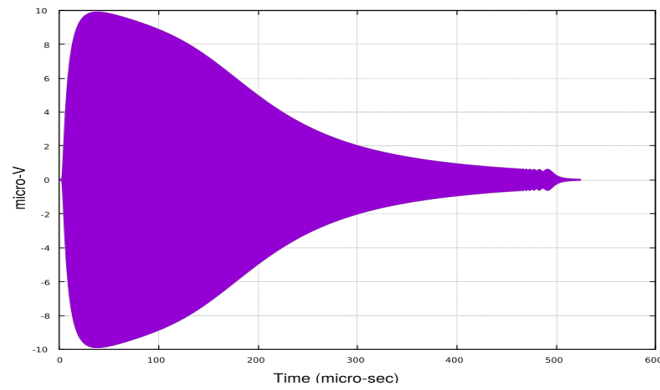
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- Can have bottom and top altitudes, or not.
- Electron density and magnetic field gradients in all directions – full description of ionospheric variations.
- Horne, JGR, vol. 94, No. A7, pp. 8895 – 8909, 1989.
  - (6) Ray trace – WKB approximation (geometric optics). Must iterate over all launch angles and frequencies.



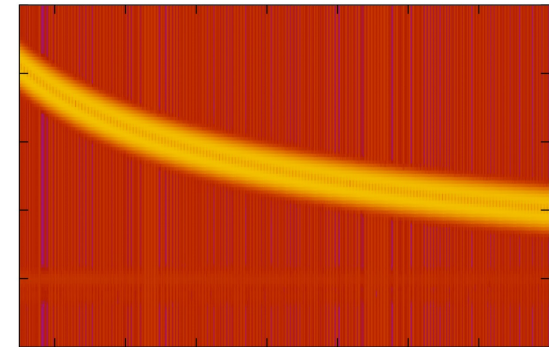
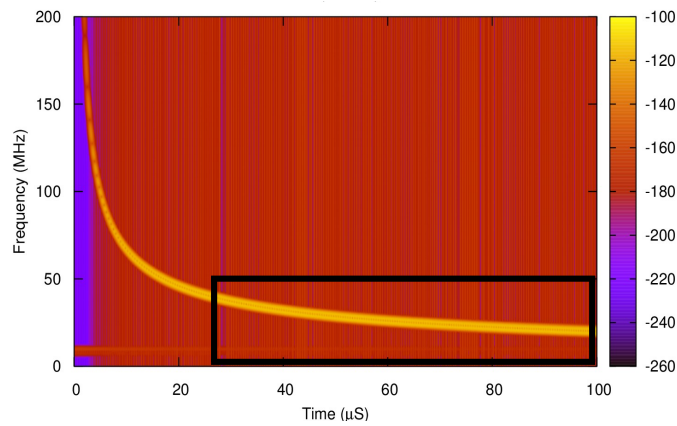
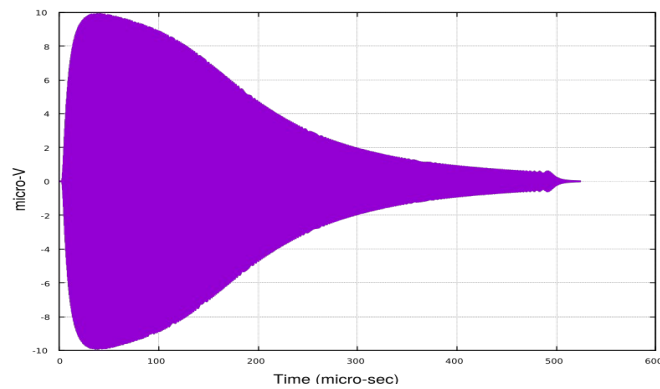
# Transionospheric Propagation

## Direct LOS - No Birefringence

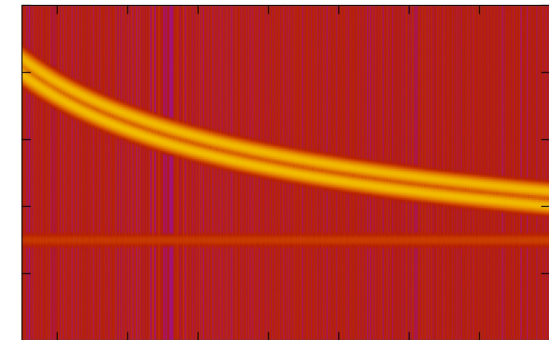
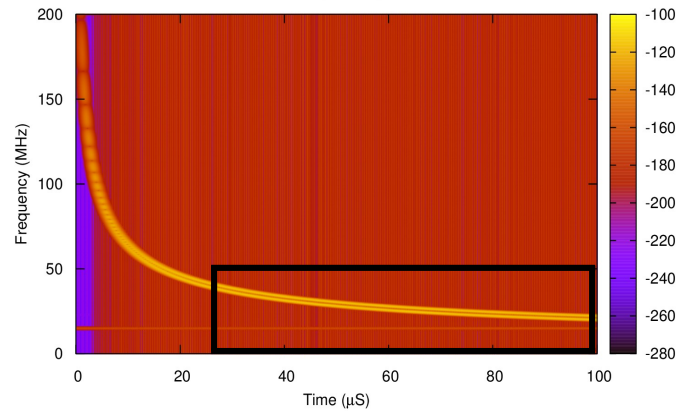
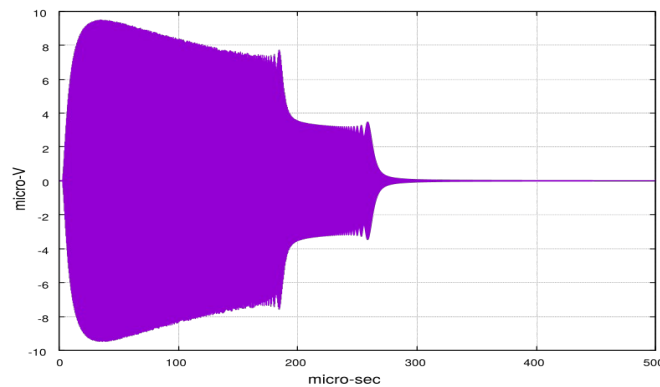


**dBm /  $\delta f$  at Satellite**

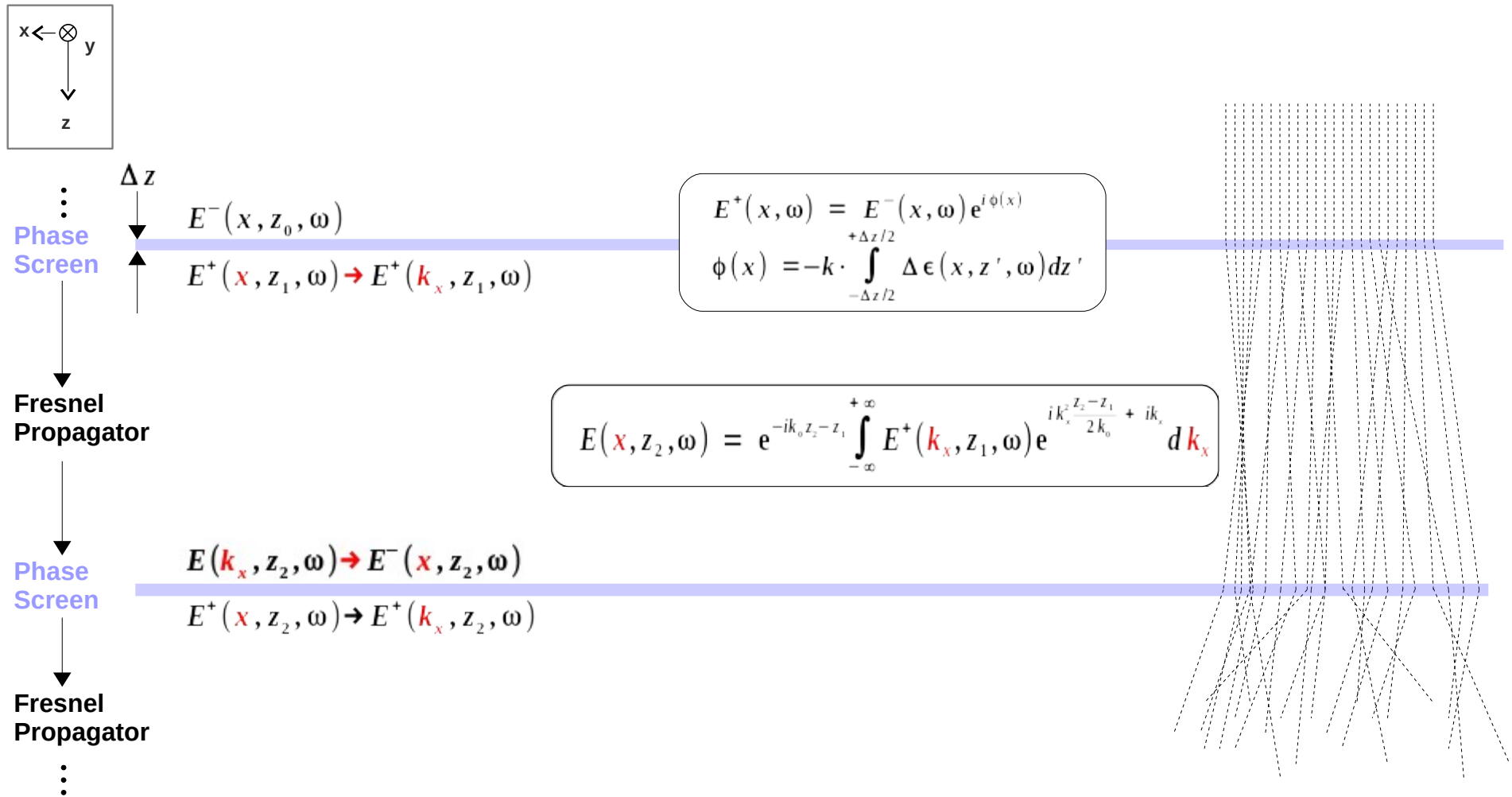
## Direct LOS - Birefringence



## Slab - 2D ray trace



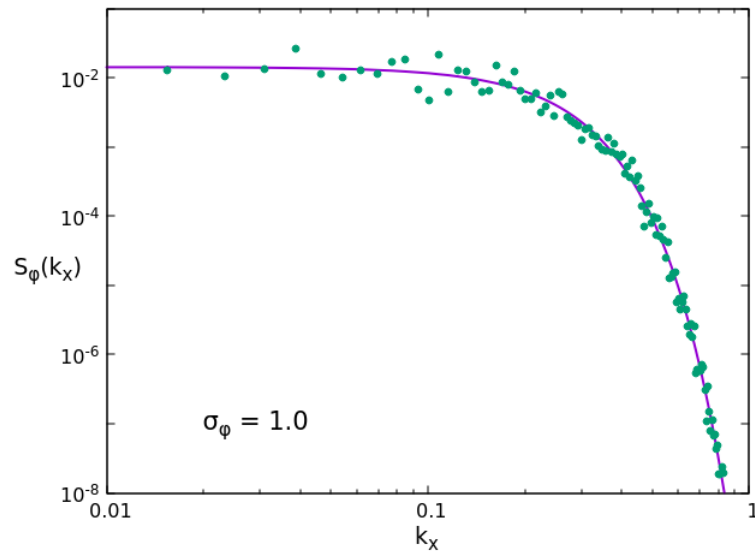
# Scintillation – Multiple Phase Screen Approach



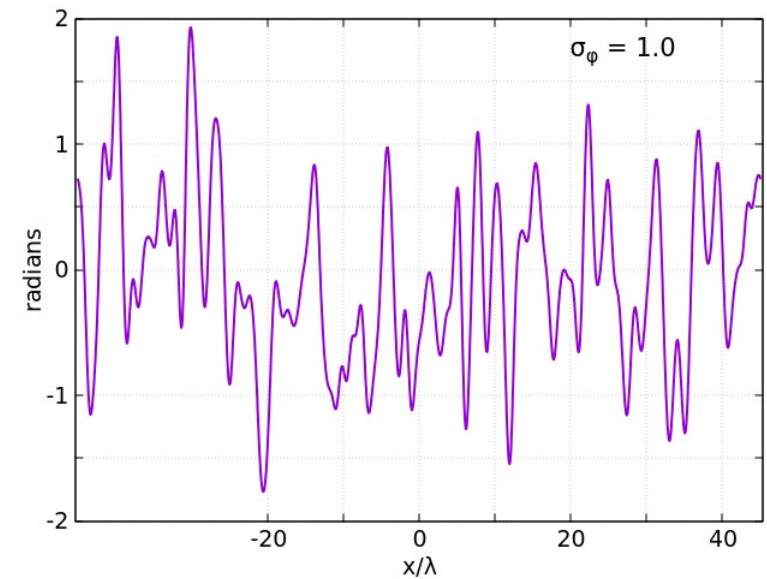
- Signal propagates through one or more thin screens
- Each screen imparts a phase contribution to the signal
- This phase is determined by the statistics of plasma density (or TEC) fluctuations
- Fresnel region can be free space or background plasma
- Implement the process by first propagating signal through a quiescent background ionosphere, and then apply phase screen(s).

# Scintillation - Phase Screen Parameters Driven by TEC Fluctuation Statistics

Statistics of TEC fluctuations



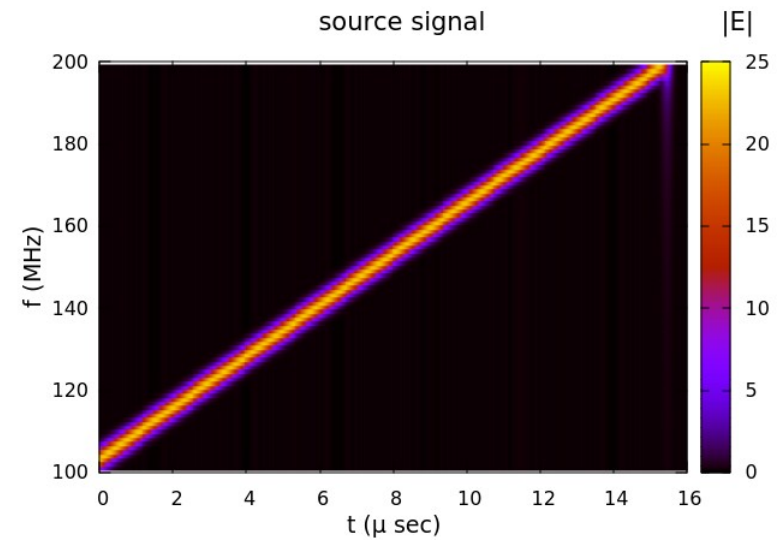
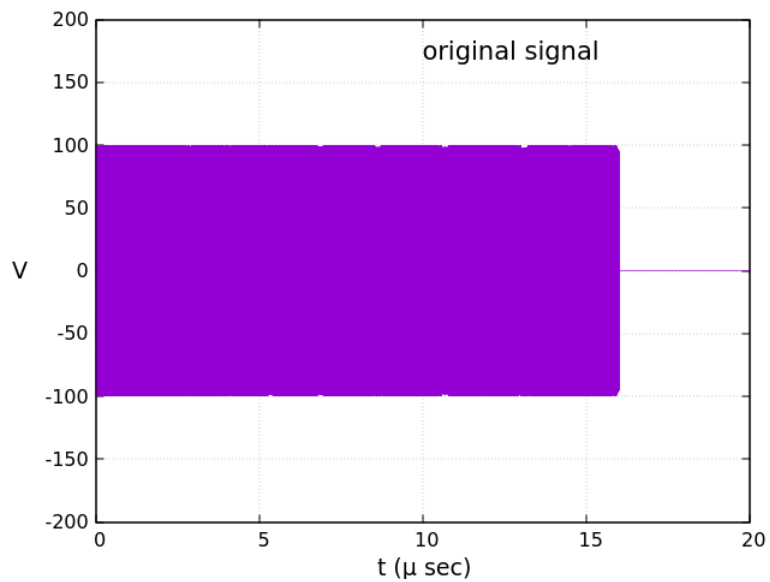
Phase Contribution  
Along Screen



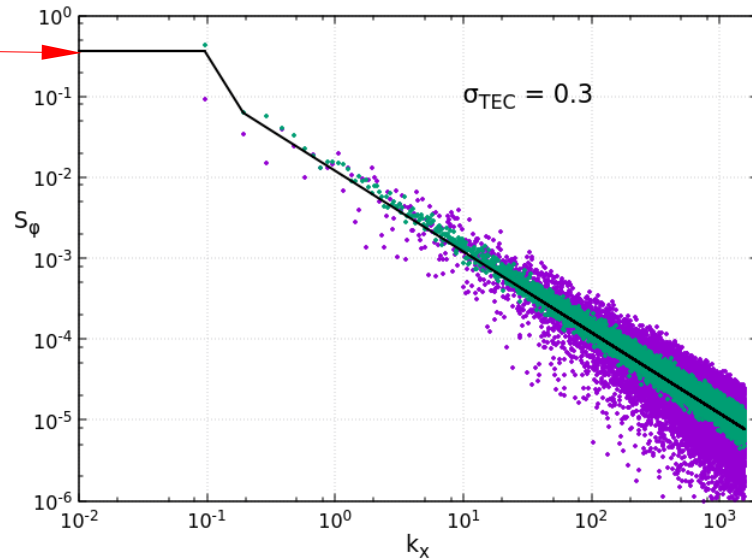
# Signal Scintillation Example

## Swept RF Pulse

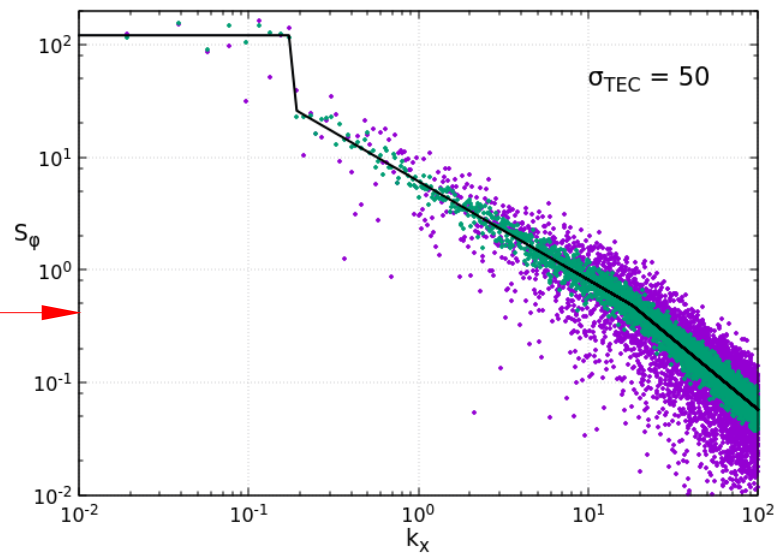
RF Pulse swept from 100 – 200 MHz



# Signal Scintillation Low and High Scintillation

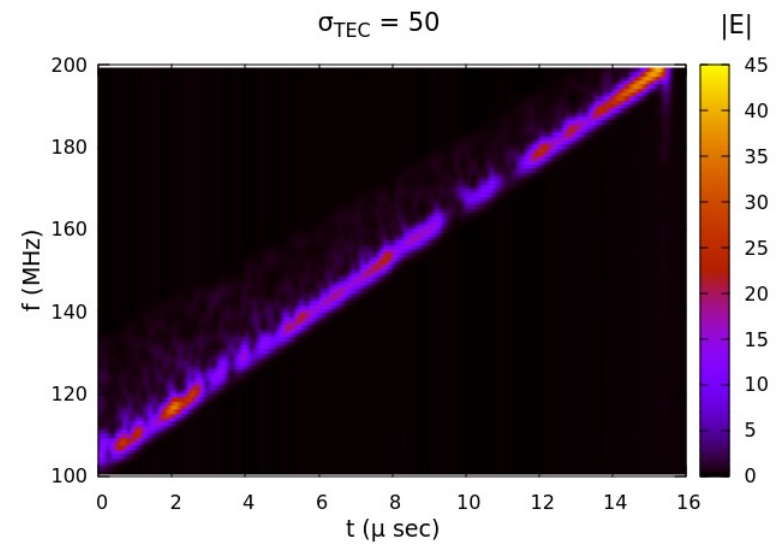
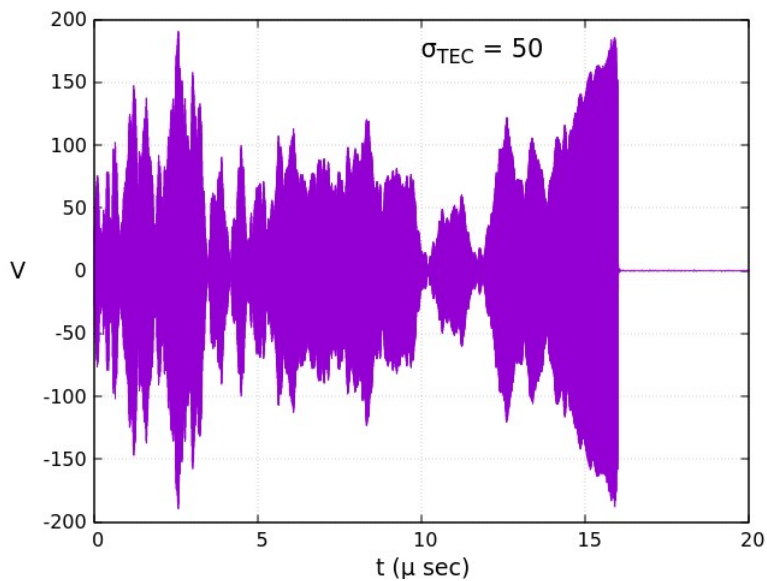
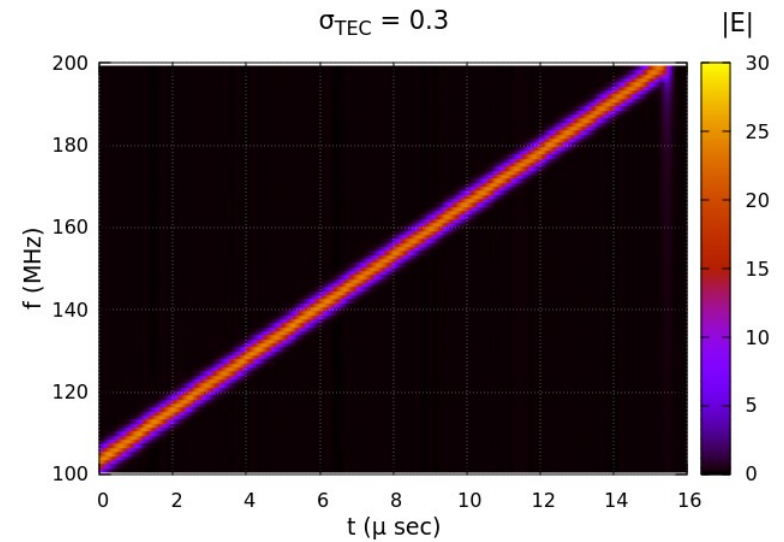
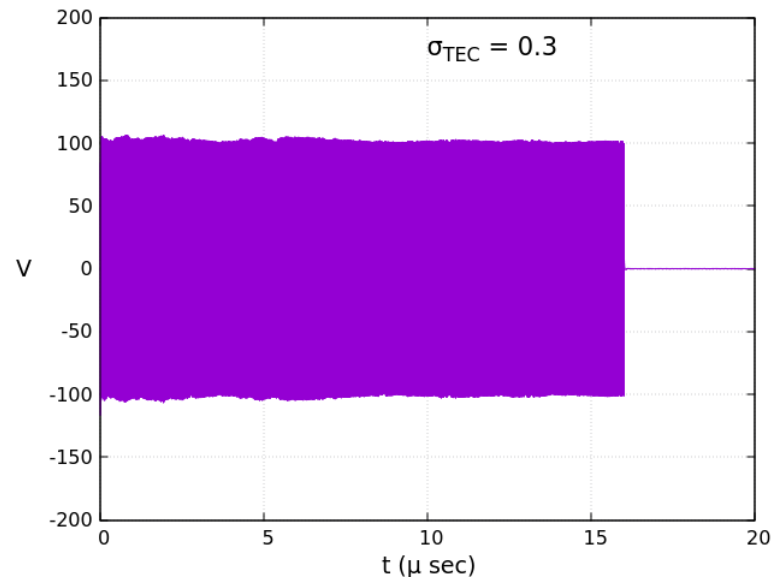


TEC variance 0.3 TECU



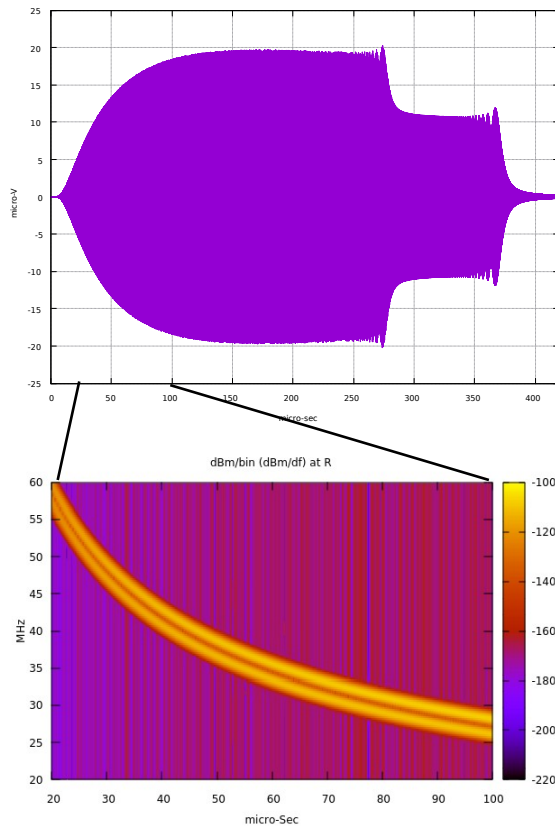
TEC variance 50 TECU

# Signal Scintillation Results

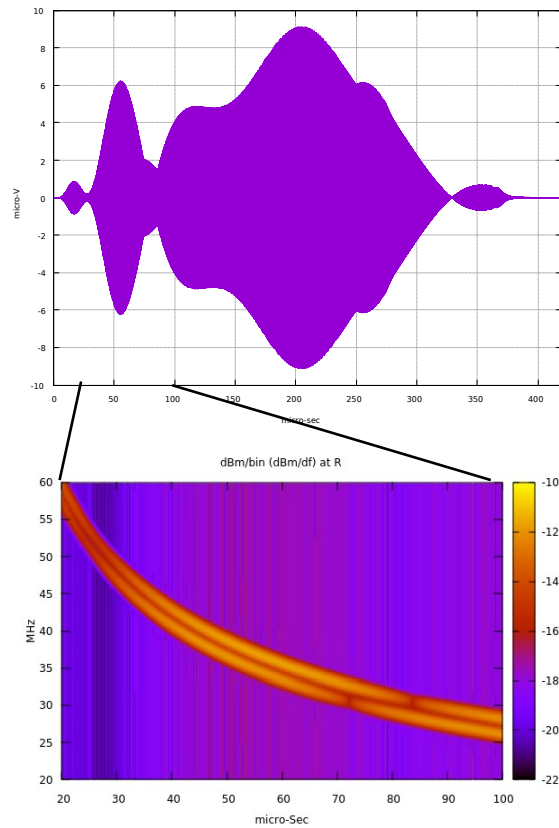


# Scintillated Transionospheric Signal using MPS Model

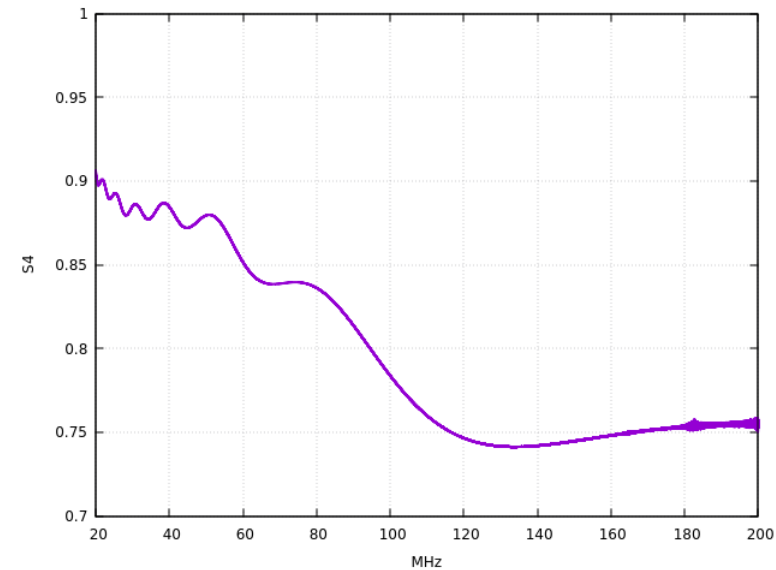
Pulse propagated through 50 TEC ionosphere using birefringent slab model.



Scintillated signal

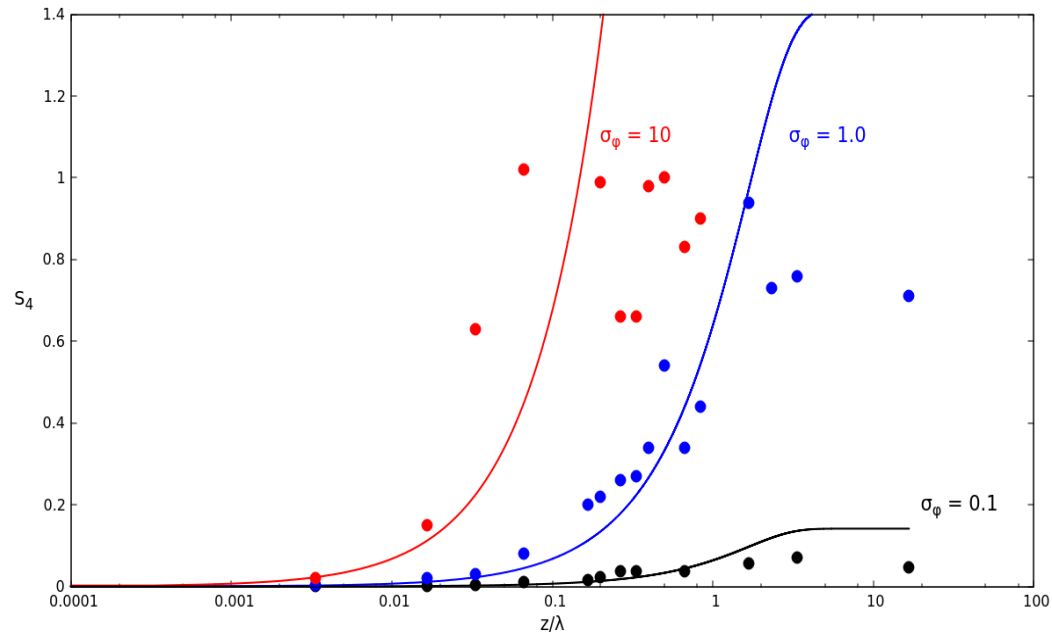


S4 over Bandwidth of Pulse

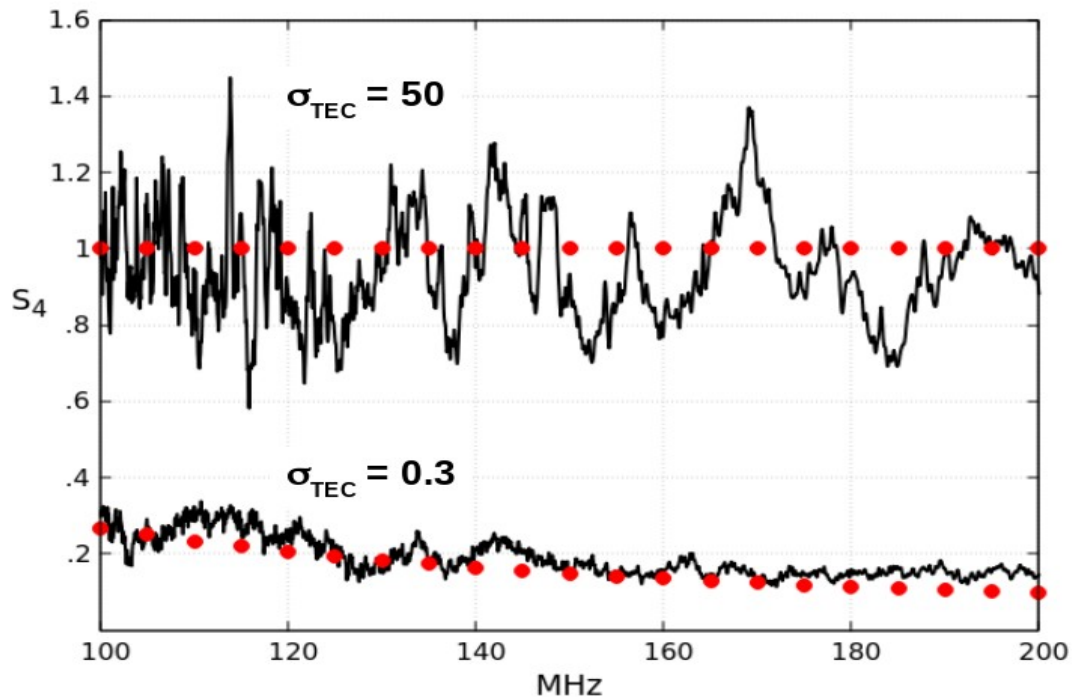


# MPS Model Results

## Comparison with Theory and Different Code



**S4 vs.  $z/\lambda$  for different  $\sigma_\phi$**   
**MPS results and weak scintillation theory**

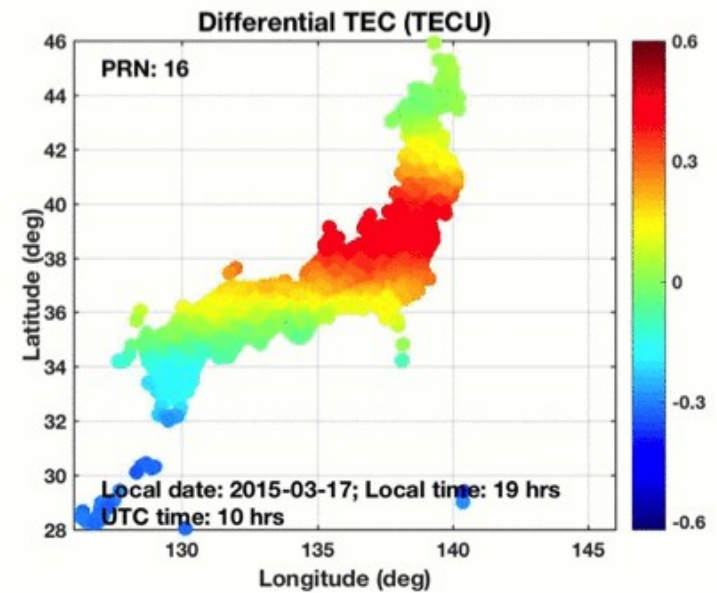
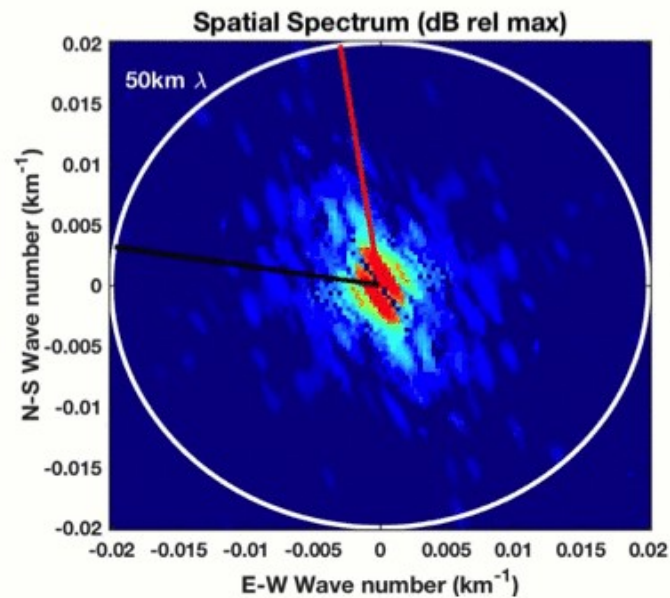


**S4 vs. frequency for different  $\sigma_\phi$**   
**MPS results and PRPSIM**



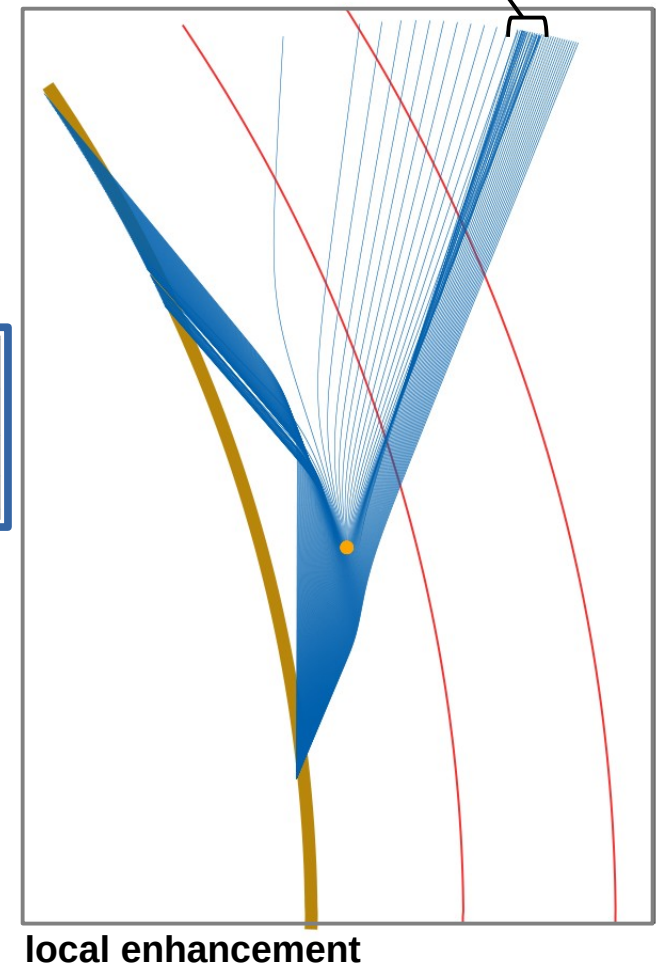
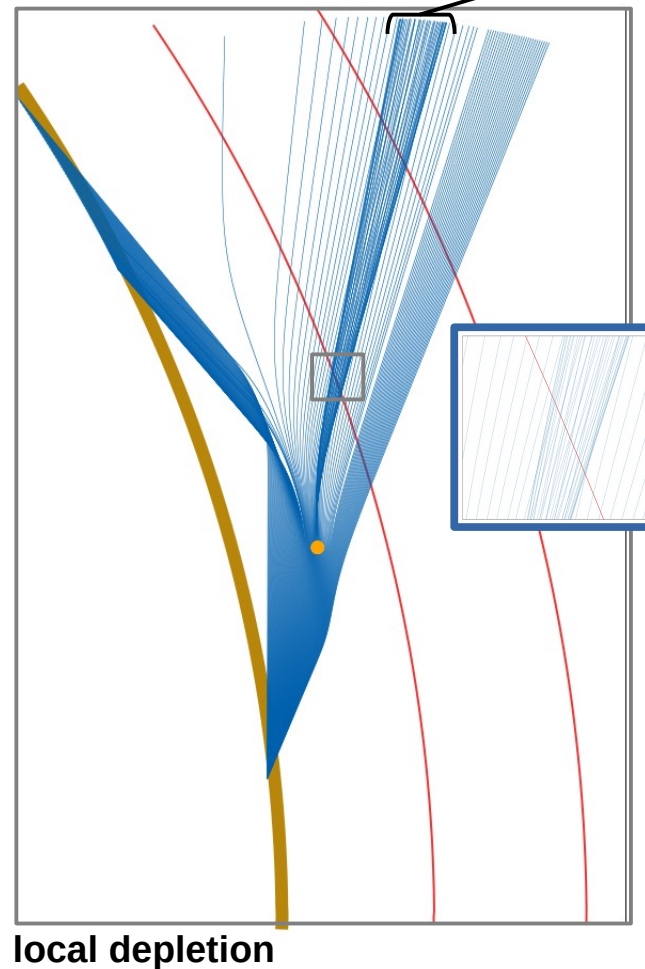
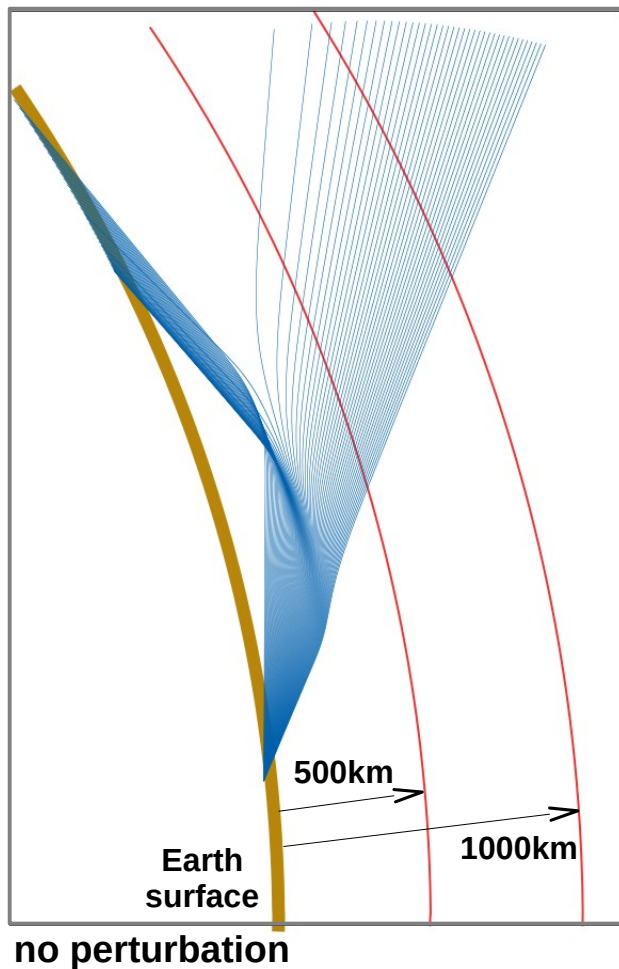
# Direct Measurement of Ionospheric disturbance k-spectrum

- 2D irregularity PSD characterization using GPS TEC receivers



## 2D, 3D Ray Trace (Doppler Spread Effect)

- Chapman density profile:  $n_{\text{max}} = 3 \times 10^{12}/\text{m}^3$ ,  $h_{\text{max}} = 300\text{km}$
- Dipole magnetic field.
- source frequency 39.7 MHz.
- source location: lat,lon =  $5^\circ, 5^\circ$ .
- depletion/enhancement location: lat,lon =  $5^\circ, 12.5^\circ$  at 300km.
- Fast Mode



Doppler spread  
regions